

# Should there be a spin-rotation coupling for a Dirac particle?

Mayeul Arminjon

*Laboratory “Soils, Solids, Structures, Risks”, 3SR  
(CNRS and Universités de Grenoble: UJF, Grenoble-INP),  
BP 53, F-38041 Grenoble cedex 9, France.*

## Abstract

It was argued by Mashhoon that a spin-rotation coupling term should add to the Hamiltonian operator in a rotating frame, as compared with the one in an inertial frame. For a Dirac particle, the Hamiltonian and energy operators were recently proved to depend on the tetrad field. We compute the energy operator in the inertial and the rotating frame, using three tetrad fields: one for each of two frameworks proposed to select the tetrad field so as to solve this non-uniqueness problem, and one proposed by Ryder. We find that Mashhoon’s term is there if the tetrad rotates as does the reference frame—but then it is also there in the Hamiltonian for the inertial frame. In fact, the Dirac Hamiltonians in two reference frames in relative rotation, but corresponding to the same tetrad field, differ only by the angular momentum term. If Mashhoon’s effect is to exist for a Dirac particle, the tetrad field must be selected in a specific way for each reference frame.

## 1 Introduction

In a reference frame that has a uniform rotation with respect to an inertial frame, the angular momentum  $\mathbf{L}$  of a particle is coupled with the rotation of the frame, in the sense that the Hamiltonian function or operator of the particle differs from its expression in the inertial frame by the term  $-\boldsymbol{\omega} \cdot \mathbf{L}$ . (Here,  $\boldsymbol{\omega}$  is the constant rotation velocity vector.) In the non-relativistic

framework (also in the presence of Newtonian gravitation), this is exact for the classical Hamiltonian function as well as for the quantum Hamiltonian operator—when, to define the latter, one considers a scalar particle without spin [1, 2]. (In a relativistic framework, for a particle without spin obeying the Klein-Gordon equation, the Hamiltonian operator in a rotating frame may have other terms involving  $\mathbf{L}$ , depending on the model metric which is considered [3, 4].) Therefore, if one considers that the spin of a quantum particle is expressing some kind of internal rotation, he may conjecture that also the spin might couple with the rotation of the reference frame. This could even be regarded [5] as a natural consequence of the fact that the total angular momentum operator is the sum of the orbital momentum and the spin. Thus, it was argued by Mashhoon [6] that a “spin-rotation coupling” term of the form

$$\delta H_{\text{SR}} = -\gamma_L \boldsymbol{\omega} \cdot \mathbf{S} \quad (1)$$

should add, in a uniformly rotating frame, to a quantum Hamiltonian  $H$  of relativistic quantum mechanics. Here,  $\gamma_L$  is the Lorentz factor corresponding to the velocity, with respect to the inertial frame, of the local observer attached with the rotating frame, and  $\mathbf{S} \equiv \frac{1}{2}\hbar\boldsymbol{\sigma}$ , where  $\boldsymbol{\sigma}$  denotes the space “vector” made with the Pauli matrices  $\sigma^j$  ( $j = 1, 2, 3$ ). The form of  $H$  was left free by Mashhoon, who got the additional term (1) from an assumption about the transformation of  $H$  and the wave function from the inertial frame to the rotating one. Later on, a similar term:

$$\delta H_{\text{SR}}' = -\boldsymbol{\omega}' \cdot \mathbf{S} \quad (2)$$

with  $\boldsymbol{\omega}'$  the “proper rotation”, was predicted by Hehl & Ni [7] to occur in the Hamiltonian of a particle obeying specifically the standard form of the (generally-)covariant Dirac equation. To write the latter explicitly, one needs to define a coordinate system and an (orthonormal) tetrad field. That prediction was obtained for a general situation in which an observer moves with a proper acceleration and a proper rotation, yet a particular tetrad field was chosen, “which behaves as a rotating Fermi-Walker-transported reference frame” [7]. Still a similar prediction was got, also from the standard covariant Dirac equation but in the case of uniform rotation, by Cai & Papini [8] who used another “rotating tetrad”. In view of these results, and since the Dirac equation is the relevant one to describe spin half particles, the spin-rotation coupling is usually considered as a theoretically well established fact. It seems that it is too small to be experimentally tested yet [2],

but it has been argued that it may have been indirectly detected [5], although the argument is not very straightforward.

Until recently, the choice of the tetrad field has been assumed to be entirely neutral, because the Lagrangian of the standard covariant Dirac equation is invariant under a change of the tetrad field, hence the Dirac equations obtained with any two different tetrad fields are equivalent [9, 10]. (This is true in a topologically-simple spacetime [11].) However, it has been observed by Ryder [12] that, in the archetypical case of uniform rotation in the Minkowski spacetime, the spin-rotation coupling term may be present or absent, depending on the choice of the tetrad field. Even more recently, it has been proved that, in a general reference frame in a general spacetime, the Hamiltonian operator  $H$  associated with the covariant Dirac equation is not unique [13]. In loose terms, the reason for this non-uniqueness is as follows: the Hamiltonian operator  $H$  is got by rewriting the wave equation in a form adapted to a particular reference frame; now, for the covariant Dirac equation, the Dirac  $\gamma^\mu$  matrices and their admissible changes are allowed to depend on the spacetime position; it follows that rewriting the covariant Dirac equation in a form adapted to a particular reference frame does not generally commute with changing the  $\gamma^\mu$  matrices in that equation. (See also Ref. [14].) It has also been proved [13] that the energy operator  $E$  is not unique, either. (The energy operator  $E$  is equal to the Hermitian part of  $H$  for the relevant scalar product and has the other important property that its mean value is the field energy [15, 16].) Nor is the spectrum of  $E$ , that is the Dirac energy spectrum. Instead, each of those three objects depends on the choice of the tetrad field, or more generally of the field of Dirac matrices [13]. Thus, contrary to a widely spread belief, the gauge invariance of the Lagrangian of the standard covariant Dirac equation—i.e., its invariance under any smooth change of the tetrad field—does not guarantee that all physically-relevant objects are also gauge-invariant. This non-uniqueness problem is there in particular in the case of an inertial frame or a uniformly rotating frame in the Minkowski spacetime [17]. That problem makes it plausible that a spin-rotation coupling term, just as the other terms of the energy operator, could be unambiguously defined only if the choice of the tetrad field were restricted in some consistent way. Note that the derivations which lead to the presence of a spin-rotation coupling term for a Dirac particle are based on choosing a tetrad that is itself rotating more or less like the reference frame [7, 8], as is also the case for Ryder’s first tetrad [12].

Whereas, Ryder's second tetrad, which does not lead to the presence of this term, is indeed non-rotating in the sense of the Fermi-Walker transport [12].

Two different frameworks have been proposed [16, 17] to restrict the choice of the tetrad field in such a way that the non-uniqueness problem [13] is proved to be solved:

**I.** With any orthonormal tetrad field  $(u_\alpha)_{\alpha=0,\dots,3}$  that is “adapted” to a given reference frame (in a sense to be precised in Sect. 3), one may associate a unique rotation rate field  $\Xi$ , which is a spatial tensor field. A first way to solve the non-uniqueness problem is to fix that spatial tensor field  $\Xi$  [17]. Two natural choices for this fixing are: *a)*  $\Xi = \Omega$ , where  $\Omega$  is the rotation-rate field of the reference frame itself [18, 19, 17]; and *b)*  $\Xi = \mathbf{0}$ . These two choices lead to non-equivalent Hamiltonians, thus represent two different solutions to the non-uniqueness problem.

**II.** A third solution is available [16] when the spacetime metric  $\mathbf{g}$  can be put in the following diagonal space-isotropic form:

$$(g_{\mu\nu}) = \text{diag}(f, -h, -h, -h), \quad f > 0, \ h > 0 \quad (3)$$

in a suitable coordinate system  $(x^\mu)$ . That other solution consists in choosing the “diagonal tetrad” in that coordinate system, i.e.,

$$u_\alpha \equiv \delta_\alpha^\mu \partial_\mu / \sqrt{|d_\mu|}, \quad d_0 \equiv f, \quad d_1 = d_2 = d_3 \equiv -h. \quad (4)$$

So the two frameworks lead to three different prescriptions for uniqueness.

The aim of this work was to compare these two frameworks for the cases of both an inertial frame and a uniformly rotating frame in the Minkowski spacetime, with special attention to the presence or absence of the Mashhoon term. Section 2 will recall the definition and the general form of the Dirac Hamiltonian and energy operators in a general spacetime. Section 3 will give some additional details about the three different prescriptions outlined above. Section 4, which contains the main new results of this paper, will apply the foregoing to the target situation. In the Minkowski spacetime, the second framework leads one to select the “Cartesian tetrad” and is very easy to put into practice. As this paper shall confirm, the first framework is much

less easy to implement. So, instead of calculating exactly the predictions of each among the two variants of the first framework, we shall determine a tetrad field which closely approaches Variant *a*). We shall also test a rotating tetrad field proposed by Ryder [12]. In each case, i.e., in the two frames and for these three tetrad fields, we shall give the explicit expression of the energy operator. We shall finally find the general expression for the difference between the Hamiltonians corresponding to a given tetrad field, in two frames having a relative rotation.

## 2 Dirac Hamiltonian and energy operators in a general spacetime

The standard form [9, 10] of the covariant Dirac equation is written in a given coordinate system  $(x^\mu)$  defined on the spacetime  $V$  (or on an open domain  $U$  therein):

$$\gamma^\mu D_\mu \Psi = -iM\Psi \quad (M \equiv mc/\hbar). \quad (5)$$

In this equation,  $\gamma^\mu$  is the field of the Dirac matrices;  $\Psi$  is the column matrix made with the components  $\Psi^a$  ( $a = 0, \dots, 3$ ) of the wave function  $\psi$ ; and  $D_\mu = \partial_\mu + \Gamma_\mu$  is the covariant derivative, where  $\Gamma_\mu$  ( $\mu = 0, \dots, 3$ ) are the connection matrices, which are  $4 \times 4$  complex matrices, just like the Dirac matrices  $\gamma^\mu$ . On the other hand,  $m$  is the rest-mass of the Dirac particles considered. The Dirac Hamiltonian operator is got by rewriting (5) in the form of the Schrödinger equation and is explicitly [20]:

$$H = mc^2\alpha^0 - i\hbar c\alpha^j D_j - i\hbar c\Gamma_0, \quad (6)$$

where

$$\alpha^0 \equiv \gamma^0/g^{00}, \quad \alpha^j \equiv \gamma^0\gamma^j/g^{00} \quad (j = 1, 2, 3). \quad (7)$$

In contrast with the wave equation (5), the Hamiltonian operator (6) changes in a non-covariant way on a general change of the coordinate system. This is true for any wave equation. However, the Hamiltonian operator transforms covariantly on a purely spatial change of the coordinates:

$$x'^0 = x^0, \quad x'^j = f^j((x^k)) \quad (j, k = 1, 2, 3). \quad (8)$$

Specifically, for the standard covariant Dirac equation, the wave function behaves as a scalar under all coordinate changes [9, 10]. It follows easily [20]

that the Dirac Hamiltonian is *invariant* after a change (8).

In the covariant Dirac equation (5), as well as in the Hamiltonian operator (6), the  $\gamma^\mu$  field is determined from the data of an orthonormal tetrad field ( $u_\alpha$ ). Decomposing the vectors  $u_\alpha$  in the natural basis ( $\partial_\mu$ ):  $u_\alpha = a^\mu{}_\alpha \partial_\mu$ , one defines

$$\gamma^\mu = a^\mu{}_\alpha \gamma^{\sharp\alpha}, \quad (9)$$

where ( $\gamma^{\sharp\alpha}$ ) is any constant “flat” set of Dirac matrices, i.e., one that is valid for the Minkowski spacetime in Cartesian coordinates [9, 10]. The definition (9) implies that the  $\gamma^\mu$  field transforms as a vector on a coordinate change (alone):

$$\gamma'^\mu = L^\mu{}_\nu \gamma^\nu, \quad L^\mu{}_\nu \equiv \frac{\partial x'^\mu}{\partial x^\nu}, \quad (10)$$

which is well known.<sup>1</sup> On the other hand, the connection matrices  $\Gamma_\mu$  transform as a covector when one changes (only) the coordinate system:

$$\Gamma'_\mu = M^\nu{}_\mu \Gamma_\nu, \quad M^\nu{}_\mu \equiv \frac{\partial x^\nu}{\partial x'^\mu}. \quad (11)$$

Using (10) and (11), the invariance of the Hamiltonian operator  $H$  under a purely spatial change (8) is also easy to check directly on the explicit form (6). We note that Eq. (11)<sub>1</sub> relates the matrices  $\Gamma_\nu$  and  $\Gamma'_\mu$  of any connection (on some vector bundle  $E$  with base  $V$ ) when two different frame fields are chosen for the tangent bundle  $TV$ , say  $(u_\nu)$  and  $(u'_\mu)$  with  $u'_\mu = M^\nu{}_\mu u_\nu$ , even if these are not coordinate bases. I.e., (11)<sub>1</sub> is true also if “non-holonomic” frame fields are chosen for  $TV$ .<sup>2</sup> This will be useful to us because, for the

---

<sup>1</sup> There are alternative versions of the covariant Dirac equation in which the wave function is a complex vector field, for which case one may optionally decompose the wave function on the coordinate basis (the natural basis of the coordinate system) [21]. Taking this option means that the frame field on the spinor bundle coincides with the coordinate basis. Then  $\Psi$  transforms as a vector and  $(\gamma^\mu)$  as a (2 1) tensor [21].

<sup>2</sup> Let  $D$  be a connection on some vector bundle  $E$  with base  $V$ , let  $(u_\alpha)$  be a frame field on  $TV$ , and let  $(e_a)$  be a frame field on  $E$ . The connection matrices  $\Gamma_\alpha$  of  $D$  in the frame fields  $(u_\alpha)$  and  $(e_a)$  are defined by their scalar components  $(\Gamma_\alpha)^b{}_a$ , such that

$$De_a(u_\alpha) = (\Gamma_\alpha)^b{}_a e_b. \quad (12)$$

This leads immediately to (11)<sub>1</sub>. If the frame field on  $TV$  is a local coordinate basis:  $u_\alpha = \delta^\mu_\alpha \partial_\mu$ , one may then compute the covariant derivatives  $D_\mu \Psi^b$  of any section of  $E$ ,  $\psi = \Psi^b e_b$ , in a matrix form:  $D_\mu \Psi = \partial_\mu \Psi + \Gamma_\mu \Psi$ . Thus, this notion of a connection

standard covariant Dirac equation, the expression of the connection matrices (of the spin connection defined on the spinor bundle) is simple if, as the frame field on TV, one chooses precisely the tetrad field  $(u_\alpha)$  used in the definition (9). These connection matrices are then [7, 12]:

$$\Gamma_\epsilon^\# = \frac{1}{8} \gamma_{\alpha\beta\epsilon} s^{\alpha\beta}, \quad s^{\alpha\beta} \equiv [\gamma^{\#\alpha}, \gamma^{\#\beta}]. \quad (13)$$

Here  $\gamma_{\alpha\beta\epsilon} \equiv \eta_{\alpha\zeta} \gamma_{\beta\epsilon}^\zeta$ , where  $\eta_{\alpha\zeta} \equiv \text{diag}(1, -1, -1, -1)$  is the Minkowski “metric” (in Cartesian coordinates) and the  $\gamma_{\beta\epsilon}^\zeta$ ’s are the coefficients of the Levi-Civita connection on TV. With an orthonormal tetrad field like  $(u_\alpha)$ , the  $\gamma_{\alpha\beta\epsilon}$ ’s can be calculated as [7, 12]:

$$\gamma_{\alpha\beta\epsilon} = -\frac{1}{2} (C_{\alpha\beta\epsilon} + C_{\beta\epsilon\alpha} - C_{\epsilon\alpha\beta}) = -\gamma_{\beta\alpha\epsilon}, \quad (14)$$

where  $C_{\alpha\beta\epsilon} \equiv \eta_{\alpha\zeta} C_{\beta\epsilon}^\zeta = -C_{\alpha\epsilon\beta}$ , the  $C_{\beta\epsilon}^\zeta$ ’s being the coefficients of the decomposition, in the tetrad basis, of the commutators of the same tetrad:

$$[u_\beta, u_\epsilon] = C_{\beta\epsilon}^\zeta u_\zeta. \quad (15)$$

**The relation (8)** between two charts (coordinate systems) is an equivalence relation for charts which are all defined on a given (open) domain U of the spacetime. We call *reference frame* an equivalence class for this relation. Thus, if  $\chi : X \mapsto (x^\mu)$  is some chart, defined on some domain U, one defines a reference frame by considering the class of  $\chi$ , that is, the set F of all charts which are defined on U and which exchange with  $\chi$  by a purely spatial change (8). A physically admissible reference frame is one for which we have  $g_{00} > 0$  everywhere in U, which condition is invariant under a change (8). The data of a physically admissible reference frame F determines [17, 18] a unique four-velocity vector  $v = v_F$ : in any chart belonging to F, its components are given by

$$(v_F)^0 \equiv \frac{1}{\sqrt{g_{00}}}, \quad (v_F)^j = 0. \quad (16)$$

---

matrix [21] extends conveniently the usual notion of the matrices of the “spin connection” entering the covariant Dirac equation, to any connection on a general vector bundle. It has a simple relation to the definition of a connection “matrix” as a matrix of one-forms [22],  $\omega = (\omega_a^b)$ : if  $(\theta^\beta)$  is the dual frame of a frame field  $(u_\alpha)$  on TV, one has  $\omega_a^b = (\Gamma_\alpha)^b_a \theta^\alpha$ . The covector transformation of the matrices  $\Gamma_\alpha$  on changing  $(u_\alpha)$  applies for a given frame field on E in (12), thus it does not apply if E = TV and  $e_a = \delta_a^\alpha u_\alpha$ .

Note that the vector  $v_F$  is indeed invariant under a change (8). Equation (16) may be rewritten as

$$v_F = \partial_0 / \sqrt{g_{00}}, \quad (17)$$

with  $(\partial_\mu)$  the natural basis of any coordinate system belonging to the frame  $F$ . This definition of a reference frame formalizes Cattaneo's idea of a reference fluid as a three-dimensional congruence of time-like world lines. The world lines of the congruence have constant space coordinates, in any chart  $\chi$  of  $F$ . The vector field  $v_F$  is the normed tangent vector field to these world lines. However, in addition, this definition fixes the time coordinate. This is necessary, because the Hamiltonian operator  $H$  does depend on the time coordinate. See Ref. [17] and references therein for more detail. Thus, the invariance of  $H$  under the changes (8) means that  $H$  depends on the coordinate system only through the reference frame. As to the dependence of  $H$  on the tetrad field, it has been proved generally in previous work [13, 17], and it will be illustrated by concrete examples in this paper. Other examples were shown in Ref. [14].

The relevant Hilbert-space scalar product was derived uniquely from rather compelling conditions [20]. This scalar product involves the hermitizing matrix  $A$  [23], which for a general  $(\gamma^\mu)$  field is also a field,  $A = A(X)$  [20]. However, usually the  $(\gamma^\mu)$  field is deduced from a tetrad field and from a constant set of “flat” Dirac matrices  $(\gamma^{\sharp\alpha})$  as in Eq. (9). Then any hermitizing matrix for the set  $(\gamma^{\sharp\alpha})$  is also a (constant) hermitizing matrix  $A$  for the  $(\gamma^\mu)$  field [20]. This is the relevant case for the present work. Moreover, in the literature, the set  $(\gamma^{\sharp\alpha})$  is usually chosen such that the hermitizing matrix is simply  $A = \gamma^{\sharp 0}$ . In that particular case, the scalar product has the form proposed by Parker [24] and by Leclerc [15]. When the operator  $H$  is not Hermitian for the scalar product (which is the general case with a non-stationary metric [15, 20, 24]), one should replace  $H$  by its Hermitian part or “energy operator”  $E$ . The latter has the physically important property that the “field energy”  $E$  associated with the Dirac field obeying Eq. (5), is equal to the mean value of the energy operator  $E$  [13, 15, 16]. However, in the present paper, only time-independent metrics and Dirac matrices  $\gamma^0$  will

occur. Therefore, the hermiticity condition proved in Ref. [20]:<sup>3</sup>

$$\partial_0 (\sqrt{-g} A \gamma^0) = 0, \quad g \equiv \det(g_{\mu\nu}) \quad (18)$$

is verified, so that *in the present paper the energy operator coincides with H*.

### 3 Different prescriptions for uniqueness

We will now give precisions about the two different frameworks [17, 16] which we proposed in order to restrict the choice of the tetrad field consistently and sufficiently, and which were outlined in Section 1. The first framework involves consideration of “spatial tensors” (e.g. “spatial vectors”), which can be defined rigorously as tensor fields on the “space manifold”  $M$  associated with a given reference frame  $F$  [17]. Here we will use simple words. As we recalled, the Hamiltonian, as well as the energy operator, depend on the reference frame. This means that, to get a unique Hamiltonian operator, we need to fix a reference frame. This can be done by considering a given physically admissible coordinate system  $(x^\mu)$  defined on some domain  $U$  of the spacetime. However, we may replace the coordinate system by another one, provided this is related to the starting one by a change (8). Let us summarize successively the two different frameworks.

**Framework I.** The data of a reference frame  $F$  fixes its four-velocity field  $v_F$ , Eq. (16). Now the vector field  $u_0$  of an orthonormal tetrad field  $(u_\alpha)$  is time-like and normed, hence it is also a four-velocity. To attach the tetrad with the reference frame, one should thus impose the condition [17, 26, 27]

$$u_0 = v_F. \quad (19)$$

Let us call this an “adapted” tetrad field to the considered reference frame  $F$ . There are many different tetrad fields which are adapted to a given arbitrary reference frame, since no condition is imposed on the vectors  $u_p$  ( $p = 1, 2, 3$ ) beyond the orthonormality of the whole tetrad  $(u_\alpha)$ . However, the latter condition implies [17] that the following tensor is antisymmetric:

$$\Phi_{\alpha\beta} \equiv g \left( u_\alpha, \left( \frac{Du_\beta}{dt} \right)_C \right) = -\Phi_{\beta\alpha}, \quad (20)$$

---

<sup>3</sup> When  $A$  is the constant  $A = \gamma^{#0}$ , the hermiticity condition has been derived in the form  $(\forall \Psi, \Phi) \int \Psi^\dagger \gamma^{#0} \partial_0 (\sqrt{-g} \gamma^0) \Phi d^3\mathbf{x} = 0$  by Parker [24] and by Huang & Parker [25]. A particular case of the latter integral condition has been derived by Leclerc [15].

where  $\left(\frac{Du}{d\xi}\right)_C$  designates the absolute derivative, with respect to the arbitrary parameter  $\xi$  along some curve  $C$  in the spacetime, of a vector  $u = u(\xi)$ . Here specifically, for any point  $X$  in the domain  $U$ , we take  $C$  to be that unique world line  $x(X)$  of the congruence attached to the reference frame  $F$  which passes at  $X$ : in any chart of  $F$ , the spatial coordinates  $x^j$  are fixed along  $x(X)$  and only the coordinate time  $t \equiv x^0/c$  varies;  $C$  is parameterized by  $t$ . We define thus  $\Phi_{\alpha\beta}(X)$ , for any point  $X \in U$ . One shows [17] that

$$\Phi_{\alpha\beta} = c \frac{d\tau}{dt} \gamma_{\alpha\beta 0}, \quad (21)$$

where  $\tau$  is the proper time along the world line  $x(X)$  and the coefficients  $\gamma_{\alpha\beta\epsilon}$  are given by Eq. (14). Moreover, one shows that the spatial components  $\Phi_{pq}$  ( $p, q = 1, 2, 3$ ) make a spatial tensor  $\Phi$  in a precise geometrical sense. This tensor is indeed the opposite of the *rotation rate of the spatial triad*  $(\mathbf{u}_p)$ . [To any four-vector  $u$ —here  $u_p$  ( $p = 1, 2, 3$ )—we associate the spatial vector  $\mathbf{u}$ —here  $\mathbf{u}_p$ —whose components are the spatial components  $u^j$  of  $u$  in a chart belonging to the reference frame  $F$  considered. This spatial vector is independent of the chart  $\chi \in F$  since, on changing the chart  $\chi \in F$  by (8), the components  $u^j$  transform correctly.] The rotation rate of the spatial triad is also a spatial tensor  $\Xi$ , whose components in the triad basis  $(\mathbf{u}_p)$  are thus:

$$\Xi_{pq} = -\Phi_{pq} = -c \frac{d\tau}{dt} \gamma_{pq0}. \quad (22)$$

It has also been proved that, if two tetrad fields are adapted to the same reference frame  $F$  and if the associated spatial triads have the same rotation rate  $\Xi$ , then the two tetrad fields give rise, in that reference frame  $F$ , to physically equivalent Dirac Hamiltonian operators, as well as to physically equivalent Dirac energy operators. Thus the first framework for uniqueness consists, in a given reference frame, in choosing an *adapted* tetrad field such that, in addition, its rotation rate tensor field  $\Xi$  is a predefined field. Two natural choices are:

- a)  $\Xi = \Omega$ , where  $\Omega$  is the rotation-rate field of the reference frame  $F$  itself [17, 18, 19], whose components in a coordinate system of  $F$  are <sup>4</sup>

$$\Omega_{jk} \equiv \frac{1}{2} c \sqrt{g_{00}} (\partial_j g_k - \partial_k g_j - g_j \partial_0 g_k + g_k \partial_0 g_j), \quad g_j \equiv \frac{g_{0j}}{g_{00}}. \quad (23)$$

---

<sup>4</sup> The spatial tensor  $\Omega$  depends on the choice of the time coordinate  $t$  in a complex manner, whereas, on changing from  $t$  to  $t'$ ,  $\Xi$  gets simply multiplied by  $dt/dt'$ . Hence, the equality  $\Xi = \Omega$  is not covariant under a change of the time coordinate, so that the

- $b) \Xi = 0$ .

As we announced in the Introduction, the two choices  $a)$  and  $b)$  lead to non-equivalent Hamiltonians, thus represent two different solutions to the non-uniqueness problem [17].

That first framework is difficult to implement, especially its variant  $a)$  which needs to calculate the field  $\Omega$  and to find a tetrad field such that  $\Xi = \Omega$ : in practice, this could be done only approximately, by numerical integration of ordinary differential equations of the form  $\delta \mathbf{u}_q/dt = \Omega^p{}_q \mathbf{u}_p$ . {Here  $\delta \mathbf{u}_q/dt$  is the Fermi-Walker derivative of  $\mathbf{u}_q$  [17].} Moreover, by imposing the condition (19), we limit from the outset the validity of this kind of solution of the non-uniqueness problem to a given reference frame.

**Framework II.** That framework needs that there is some special coordinate system  $(x^\mu)$ , in which the metric has the special form (3) [16]. As discussed there, this form is general enough for the prospective purpose of testing the generally-covariant Dirac equations in a realistic spacetime metric.<sup>5</sup> Then one chooses the “diagonal tetrad” in that coordinate system, Eq. (4). This defines the Dirac matrices  $\gamma^\mu$  in that coordinate system, Eq. (9). Then, in any possible coordinate system, say  $(x'^\mu)$ , the Dirac matrices  $\gamma'^\mu$  are got by the transformation (10). As it has been proved in Ref. [16]: if one considers another coordinate system in which the metric has also the *form* (3) (a priori not with the same coefficients), then one passes from the first to the second one by a constant rotation, combined with a constant homothecy.

---

prescriptions  $\Xi = \Omega$  corresponding to reference frames differing merely in the choice of the time coordinate are not physically equivalent. And indeed, there is a rewriting of the geodesic equation of motion in the form of Newton’s second law, in which the tensor  $\Omega$  plays exactly the role played by the angular velocity tensor of a rotating frame in Newtonian theory [18]—but in this rewriting  $\Omega$  has to be calculated with a time coordinate  $\hat{x}^0$  such that, along a world line of the congruence, we have  $d\hat{x}^0 = c d\tau$ , where  $d\tau$  is the proper time increment. Thus, if one applies the prescription  $\Xi = \Omega$ , one should impose that the time coordinate be a such one,  $\hat{x}^0$  with  $d\hat{x}^0 = c d\tau$  [17]. For the uniformly rotating frame, the tensors  $\Omega$  calculated with either  $t$  or  $\tau$  differ only by  $O(V^2/c^2)$  [17] ( $V$  is defined in Sect. 4), and the same is easy to check for  $\Xi$ .

<sup>5</sup> Moreover, this form is generic for an alternative theory of gravitation [28], in the preferred reference frame assumed by that theory. That theory is based only on a scalar field which determines, among other things, the physical metric  $\mathbf{g}$ , from an a priori assumed flat metric, say  $\gamma$ . Although it thus has two metrics, this is not a metric theory in the standard sense.

It follows [16], first, that the corresponding “diagonal tetrads” (4) exchange by a *constant* Lorentz transformation, and then, that in any given reference frame, the Hamiltonian operators got from these two choices of tetrad fields are equivalent, as well as the energy operators. Thus the non-uniqueness problem is solved simultaneously in any possible reference frame, and in a simple tractable way.

## 4 Dirac energy operator in an inertial or a rotating frame

Starting with an inertial reference frame  $F'$  in a Minkowski spacetime, defined from a Cartesian system of coordinates  $(x'^\mu) = (ct', x', y', z')$ , we define the uniformly rotating reference frame  $F$  from the rotating coordinates  $(x^\mu) = (ct, x, y, z)$  given by

$$t = t', \quad x = x' \cos \omega t + y' \sin \omega t, \quad y = -x' \sin \omega t + y' \cos \omega t, \quad z = z', \quad (24)$$

where  $\omega$  is a real constant. In these coordinates, the Minkowski metric remains stationary: it becomes

$$ds^2 = \left[ 1 - \left( \frac{\omega}{c} \right)^2 (x^2 + y^2) \right] (dx^0)^2 + 2 \frac{\omega}{c} (y dx - x dy) dx^0 - (dx^2 + dy^2 + dz^2). \quad (25)$$

The validity of these new coordinates is restricted by the admissibility condition  $g_{00} > 0$  to the domain  $U$  made of those points in the spacetime for which we have  $V \equiv \omega \rho < c$ , where  $\rho \equiv (x^2 + y^2)^{1/2}$ .

### 4.1 Energy operators in the two frames with the Cartesian tetrad

In the Cartesian coordinates  $(x'^\mu)$  on the Minkowski spacetime, the metric has of course the space-isotropic diagonal form (3), hence we can apply Framework II. The corresponding diagonal tetrad (4) is just  $u'_\alpha \equiv \delta'_\alpha{}^\mu \partial'_\mu$ , that is, the natural basis of the Cartesian coordinate system  $(x'^\mu)$ , or “Cartesian tetrad”. Clearly, the coefficients  $\gamma^\zeta_{\beta\epsilon}$  of the Levi-Civita connection are zero

with this tetrad field  $(u'_\alpha)$ ,<sup>6</sup> so the connection matrices (13) are  $\Gamma^\sharp_\epsilon = 0$  and, by (11), they remain zero in any coordinates. Also, using the tetrad  $(\partial'_\mu)$ , the Dirac matrices (9) in the coordinates  $(x'^\mu)$  are simply the “flat” matrices,  $\gamma'^\mu = \gamma^{\sharp\mu}$ . Hence, when it is used in the inertial frame  $F'$  itself, the Cartesian tetrad yields by (6) just the special-relativistic Hamiltonian, which is Hermitian:

$$E'_1 = H'_1 = mc^2\gamma^{\sharp 0} - i\hbar c\alpha^{\sharp j}\partial'_j, \quad (26)$$

where  $\alpha^{\sharp j} \equiv \gamma^{\sharp 0}\gamma^{\sharp j}$ . This result is the physically correct one: in an inertial frame, the Hamiltonian operator should indeed be the one predicted by Dirac’s original theory.

The Hamiltonian  $H_1$  in the rotating frame  $F$  and corresponding with the Cartesian tetrad involves the Dirac matrices transformed to the rotating coordinates (24) by Eq. (10):

$$\gamma^0 = \gamma^{\sharp 0}, \quad \gamma^1 = \gamma^{\sharp 1} \cos \omega t + \gamma^{\sharp 2} \sin \omega t + \frac{\omega y}{c} \gamma^{\sharp 0}, \quad (27)$$

$$\gamma^2 = -\gamma^{\sharp 1} \sin \omega t + \gamma^{\sharp 2} \cos \omega t - \frac{\omega x}{c} \gamma^{\sharp 0}, \quad \gamma^3 = \gamma^{\sharp 3}. \quad (28)$$

From (18), (25) and (27)<sub>1</sub>, it follows that  $H_1$  is Hermitian. Noting that  $g^{00} = 1$  after the coordinate change (24), we get then the  $\alpha$  matrices of Eq. (7):

$$\alpha^0 = \gamma^{\sharp 0}, \quad \alpha^1 = \alpha^{\sharp 1} \cos \omega t + \alpha^{\sharp 2} \sin \omega t + \frac{\omega y}{c} \mathbf{1}_4, \quad (29)$$

$$\alpha^2 = -\alpha^{\sharp 1} \sin \omega t + \alpha^{\sharp 2} \cos \omega t - \frac{\omega x}{c} \mathbf{1}_4, \quad \alpha^3 = \alpha^{\sharp 3}. \quad (30)$$

We have moreover from (24):

$$\cos \omega t \partial_x - \sin \omega t \partial_y = \partial_{x'}, \quad \sin \omega t \partial_x + \cos \omega t \partial_y = \partial_{y'}, \quad \partial_z = \partial_{z'}. \quad (31)$$

Therefore, the energy operator  $E_1 = H_1$ , Eq. (6), is explicitly:

$$\begin{aligned} H_1 &= mc^2\alpha^0 - i\hbar c\alpha^j\partial_j \\ &= mc^2\gamma^{\sharp 0} - i\hbar c \left[ \alpha^{\sharp 1}(\cos \omega t \partial_x - \sin \omega t \partial_y) + \alpha^{\sharp 2}(\sin \omega t \partial_x + \cos \omega t \partial_y) + \alpha^{\sharp 3}\partial_z \right] \\ &\quad - i\hbar c \frac{\omega}{c} (y\partial_x - x\partial_y) \mathbf{1}_4 \\ &= H'_1 - i\hbar\omega(y\partial_x - x\partial_y), \\ H_1 &= H'_1 - \boldsymbol{\omega} \cdot \mathbf{L}. \end{aligned} \quad (32)$$

---

<sup>6</sup> Hence, by (22): for that tetrad, in the inertial frame  $F'$  to which it is adapted, we have  $\Xi = \mathbf{0}$ : the Cartesian tetrad solves Variant b) of Framework I for the inertial frame.

Here,  $\mathbf{L} \equiv \mathbf{r} \wedge (-i\hbar\nabla)$  is the angular momentum operator. Thus, in the case of a uniformly rotating frame in a Minkowski spacetime [and arguably in general, see Eq. (75) below], Framework II does not predict any spin-rotation coupling.

## 4.2 Constructing a tetrad adapted to the rotating frame

Let us now try to use Framework I. One rotating orthonormal tetrad that appears naturally in the metric (25) is Ryder's [12] first tetrad:

$$u_0 = \frac{1}{c} \frac{\partial}{\partial t} + \frac{\omega y}{c} \frac{\partial}{\partial x} - \frac{\omega x}{c} \frac{\partial}{\partial y}, \quad u_1 = \frac{\partial}{\partial x}, \quad u_2 = \frac{\partial}{\partial y}, \quad u_3 = \frac{\partial}{\partial z}. \quad (33)$$

However, as noted in Ref. [17], it results from (24) and (33) that

$$u_0 = \frac{\partial x^\nu}{\partial x'^0} \frac{\partial}{\partial x^\nu} = \frac{\partial}{\partial x'^0} \equiv \partial'_0, \quad (34)$$

$$u_1 = \cos \omega t \partial'_1 + \sin \omega t \partial'_2, \quad u_2 = -\sin \omega t \partial'_1 + \cos \omega t \partial'_2, \quad u_3 = \partial'_3, \quad (35)$$

where  $(\partial'_\mu)$  is the Cartesian tetrad. Since  $v_F = \partial_0/\sqrt{g_{00}}$ , Eq. (17), Equation (34) means that Ryder's tetrad  $(u_\alpha)$  is “adapted” in the sense of Eq. (19) to the inertial frame  $F'$ , not to the rotating frame  $F$ . On the other hand, since  $\mathbf{g}(\partial_\mu, \partial_\nu) = g_{\mu\nu}$ , we see from (25) that the natural basis  $(\partial_\mu)$  of the rotating coordinates  $x^\mu$  given by (24) is not orthogonal. But consider, at each  $X \in U$ , the hyperplane  $H_X$  in the local tangent space  $TV_X$  to the spacetime  $V$ , made of the vectors which are orthogonal to  $v_F(X)$ . Define the orthogonal projection  $\Pi_X$  onto  $H_X$  [17, 29]. (Note that this operator depends on the reference frame  $F$  which is considered, as do  $v_F$  and  $H_X$ .) Obviously, if at each  $X \in U$  we thus project the vectors  $\partial_j(X)$  ( $j = 1, 2, 3$ ) onto  $H_X$ , we get vector fields  $\Pi\partial_j$  such that  $(\Pi\partial_j)(X)$  is orthogonal to  $v_F(X)$  at any  $X \in U$ . From the definition, one finds the components of  $\Pi_X a$  for a vector  $a \in TV_X$ , in a coordinate system belonging to  $F$  [17]. Thus we get

$$(\Pi\partial_j)^0 = -g_{0k}(\partial_j)^k/g_{00} = -g_{0j}/g_{00}, \quad (\Pi\partial_j)^k = (\partial_j)^k = \delta_j^k, \quad (36)$$

from which it follows that

$$\mathbf{g}(\Pi\partial_j, \Pi\partial_k) = g_{jk} - \frac{g_{0j}g_{0k}}{g_{00}} \equiv -h_{jk}. \quad (37)$$

Here  $\mathbf{h}$  is the spatial metric of the reference frame F, such that for any two vectors  $a, b$  at  $X$  [17, 29]:

$$\mathbf{h}_X(a, b) \equiv -\mathbf{g}_X(\Pi_X a, \Pi_X b). \quad (38)$$

Note that the definition of  $\mathbf{H}_X$  and  $\Pi_X$ , as well as Eqs. (36) to (38), are valid for a general reference frame in a general spacetime. Coming back to the uniformly rotating frame in a Minkowski spacetime, from (25) and (37) we get that  $\mathbf{g}(\Pi\partial_1, \Pi\partial_3) = \mathbf{g}(\Pi\partial_2, \Pi\partial_3) = 0$  but  $\mathbf{g}(\Pi\partial_1, \Pi\partial_2) \neq 0$ ,  $(\partial_\mu)$  being again specifically the natural basis associated with the coordinates (24). Hence, one may define an orthonormal tetrad adapted to the rotating frame F by taking:  $v_F, \Pi\partial_2 / \|\Pi\partial_2\|, \Pi\partial_3 / \|\Pi\partial_3\|$ , and the vector product of (the spatial vectors associated with) the two last vectors. However, a simpler orthonormal tetrad adapted to F is obtained by considering the natural basis  $(\partial_\mu^\circ)$  of the “rotating cylindrical coordinates”  $(x^\circ) = (ct, \rho, \varphi, z)$ , related to the coordinates (24) by

$$x = \rho \cos \varphi, \quad y = \rho \sin \varphi. \quad (39)$$

(It follows from this that  $\partial_0^\circ = \partial_0$  and  $\partial_3^\circ = \partial_z^\circ = \partial_3 = \partial_z$ .) In the coordinates  $(x^\circ)$ , the Minkowski metric (25) rewrites immediately as

$$ds^2 = g_{\mu\nu}^\circ dx^{\circ\mu} dx^{\circ\nu} = \left[ 1 - \left( \frac{\omega\rho}{c} \right)^2 \right] c^2 dt^2 - 2\omega\rho^2 d\varphi dt - (d\rho^2 + \rho^2 d\varphi^2 + dz^2), \quad (40)$$

from which we find that the spatial metric defined by Eq. (37)<sub>2</sub> has components [in the coordinates  $(x^\circ) = (\rho, \varphi, z)$ ]:

$$h_{jk} = \delta_{jk} \quad \text{except for} \quad h_{22} = \frac{\rho^2}{1 - \omega^2 \rho^2 / c^2}. \quad (41)$$

Hence, owing to Eq. (37)<sub>1</sub>, we define an orthonormal tetrad adapted to the rotating frame F by taking  $v_F = \partial_0^\circ / \sqrt{g_{00}^\circ}$  and by norming the  $\Pi\partial_j^\circ$  vectors, which results simply in setting

$$u_0^\circ = \frac{1}{\sqrt{1 - \omega^2 \rho^2 / c^2}} \partial_0, \quad u_1^\circ = \partial_1^\circ = \partial_\rho, \quad (42)$$

$$u_2^\circ = \frac{\omega\rho}{c\sqrt{1 - \omega^2 \rho^2 / c^2}} \partial_0 + \frac{\sqrt{1 - \omega^2 \rho^2 / c^2}}{\rho} \partial_\varphi, \quad u_3^\circ = \partial_3^\circ = \partial_z^\circ = \partial_z. \quad (43)$$

We note that the matrix  $a \equiv (a^\mu_\alpha)$ , such that  $u^\circ_\alpha = a^\mu_\alpha \partial^\circ_\mu$ , is independent of the time coordinate  $t$ . Hence, so are also the Dirac matrices (9). Thus, from (18), the Hamiltonian operator in the rotating frame with the adapted rotating tetrad is Hermitian.

Let us calculate the rotation rate tensor field  $\Xi$  of the tetrad  $(u^\circ_\alpha)$ , Eq. (22). The coefficients of the decomposition (15) of the commutators of the tetrad (42)-(43) are easily computed to be:  $C^\zeta_{\beta\epsilon} = 0$ , except for:

$$C^0_{01} = -C^0_{10} = -\frac{\rho\omega^2}{c^2 - \omega^2\rho^2}, \quad C^0_{12} = -C^0_{21} = \frac{2\omega}{c(1 - \omega^2\rho^2/c^2)}, \quad (44)$$

$$C^2_{12} = -C^2_{21} = -\frac{1}{\rho(1 - \omega^2\rho^2/c^2)}. \quad (45)$$

From this, we deduce immediately the coefficients  $C_{\alpha\beta\epsilon} = \eta_{\alpha\zeta} C^\zeta_{\beta\epsilon}$ , then we get the coefficients  $\gamma_{\alpha\beta\epsilon} = -\gamma_{\beta\alpha\epsilon}$  [Eq. (14)]. They are zero, except for (when  $\alpha < \beta$ ):

$$\gamma_{010} = -\frac{\rho\omega^2}{c^2 - \omega^2\rho^2}, \quad \gamma_{122} = \frac{1}{\rho(1 - \omega^2\rho^2/c^2)}, \quad (46)$$

$$\gamma_{120} = -\gamma_{012} = \gamma_{021} = \frac{\omega}{c(1 - \omega^2\rho^2/c^2)}. \quad (47)$$

Therefore, Eqs. (22) and (40) give us:  $\Xi_{pq} = 0$ , except for

$$\Xi_{21} = -\Xi_{12} = \omega\gamma_L, \quad \gamma_L = \gamma_L(\rho) \equiv [1 - (\omega^2\rho^2/c^2)]^{-1/2}. \quad (48)$$

We may compare this with the rotation rate tensor  $\Omega$  of the reference frame, defined in general by Eq. (23). For the rotating frame F, the components  $\Omega_{jk}$  of  $\Omega$  are easily computed [17]:

$$\Omega_{32} = 0, \quad \Omega_{13} = 0, \quad \Omega_{21} = +\omega\gamma_L^3. \quad (49)$$

These are in fact the components of the spatial tensor  $\Omega$  in the natural basis  $(\partial_j)$  associated with the spatial coordinates  $(x^j)$  [the spatial part of the coordinates (24)]. The components  $\Omega^\circ_{pq}$  of  $\Omega$  in the spatial triad basis  $(\mathbf{u}^\circ_p)$  associated with the tetrad basis  $(u^\circ_\alpha)$  are got from (49) and from the relation between the triad bases  $(\partial_j)$  and  $(\mathbf{u}^\circ_p)$ . This relation follows from (39) and (42)-(43) and is:

$$\mathbf{u}^\circ_1 = \cos \varphi \partial_1 + \sin \varphi \partial_2, \quad \mathbf{u}^\circ_2 = (-\sin \varphi \partial_1 + \cos \varphi \partial_2) / \gamma_L(\rho), \quad \mathbf{u}^\circ_3 = \partial_3. \quad (50)$$

By standard tensor transformation, we find from this and from (49):

$$\Omega_{32}^\circ = 0, \quad \Omega_{13}^\circ = 0, \quad \Omega_{21}^\circ = \Omega_{21}/\gamma_L = \omega\gamma_L^2. \quad (51)$$

These differ from (48) only by  $O(V^2/c^2)$  terms (for  $V \equiv \omega\rho \ll c$ ). Up to this negligible difference, we may thus consider that the adapted rotating tetrad  $(u_\alpha^\circ)$  verifies  $\Xi = \Omega$ , as required by the variant *a*) of Framework I.

### 4.3 Energy operator with the adapted rotating tetrad

Let us thus calculate the Hamiltonian (6) in the rotating frame F, when choosing the tetrad  $(u_\alpha^\circ)$ , Eqs. (42)-(43). We begin with the spin connection matrices (13) with the tetrad field  $(u_\alpha^\circ)$ . From Eqs. (46)-(47), these are:

$$4\Gamma_0^\sharp = \gamma_{010} s^{01} + \gamma_{120} s^{12}, \quad 4\Gamma_1^\sharp = \gamma_{021} s^{02}, \quad (52)$$

$$4\Gamma_2^\sharp = \gamma_{012} s^{01} + \gamma_{122} s^{12}, \quad \Gamma_3^\sharp = 0. \quad (53)$$

To compute the connection matrices  $\Gamma_\mu$  when the coordinate basis  $(\partial_\mu^\circ)$  is chosen, we use the fact that they transform as a covector [see Eq. (11) and thereafter]. Thus we have  $\Gamma_\mu = b^\alpha{}_\mu \Gamma_\alpha^\sharp$ , where the matrix  $b \equiv (b^\alpha{}_\mu)$ , such that  $\partial_\mu^\circ = b^\alpha{}_\mu u_\alpha^\circ$ , is got easily from Eqs. (42)-(43):

$$b = \begin{pmatrix} \gamma_L^{-1} & 0 & -\frac{\gamma_L \omega \rho^2}{c} & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \rho \gamma_L & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}. \quad (54)$$

We get thus from (52)-(53), using the standard set of Dirac matrices: <sup>7</sup>

$$\Gamma_0 = -\frac{\gamma_L}{2} \left( \frac{i\omega}{c} \Sigma^3 + \frac{\rho\omega^2}{c^2} \Sigma'^1 \right), \quad \Gamma_1 = \frac{i\omega\gamma_L^2}{2c} \Sigma'^2, \quad (55)$$

$$\Gamma_2 = \frac{\gamma_L^3}{2} \left[ \left( \frac{\rho\omega}{c} \right)^2 - 1 \right] \left( \frac{\rho\omega}{c} \Sigma'^1 + i\Sigma^3 \right), \quad \Gamma_3 = 0, \quad (56)$$

---

<sup>7</sup> The choice of the set  $(\gamma^{\sharp\alpha})$  does not matter, because corresponding  $(\gamma^\mu)$  fields exchange by constant similarity transformations, hence give rise to equivalent energy operators. With the standard set (Dirac's), we have  $s^{jk} = -2i\epsilon_{jkl}\Sigma^l$  and  $s^{0j} = 2\Sigma'^j$ .

where

$$\Sigma^j \equiv \begin{pmatrix} \sigma^j & 0 \\ 0 & \sigma^j \end{pmatrix}, \quad \Sigma'^j \equiv \begin{pmatrix} 0 & \sigma^j \\ \sigma^j & 0 \end{pmatrix}. \quad (57)$$

On the other hand, the  $\gamma^\mu$  matrices are defined by (9). In view of Eqs. (42)-(43), we have:

$$\gamma^0 = \gamma_L \gamma^{\#0} + \frac{\rho \gamma_L \omega}{c} \gamma^{\#2}, \quad \gamma^1 = \gamma^{\#1}, \quad \gamma^2 = \frac{1}{\rho \gamma_L} \gamma^{\#2}, \quad \gamma^3 = \gamma^{\#3}, \quad (58)$$

from which we get the matrices  $\alpha^\mu$  of Eq. (7) [note that  $g^{00} = 1$ ]:

$$\alpha^0 = \gamma^0, \quad \alpha^2 = \frac{1}{\rho} \left( \alpha^{\#2} + \frac{\rho \omega}{c} \mathbf{1}_4 \right) = \frac{1}{\rho} \left( \Sigma'^2 + \frac{\rho \omega}{c} \mathbf{1}_4 \right), \quad (59)$$

$$\alpha^j = \gamma_L \left( \alpha^{\#j} + \frac{\rho \omega}{2c} s^{2j} \right) = \gamma_L \left( \Sigma'^j - i \frac{\rho \omega}{c} \epsilon_{2jk} \Sigma^k \right) \quad (j = 1, 3). \quad (60)$$

The energy operator with the adapted rotating tetrad is thus [Eq. (6)]:

$$E_2 = H_2 = mc^2 \alpha^0 - i \hbar c \alpha^j (\partial_j^\circ + \Gamma_j) - i \hbar c \Gamma_0, \quad (61)$$

where the matrices  $\Gamma_\mu$  and  $\alpha^\mu$  are given by Eqs. (55)-(56) and (59)-(60). In particular, for  $V \equiv \rho \omega \ll c$ , we have from (55):

$$-i \hbar c \Gamma_0 = -\frac{\hbar \gamma_L \omega}{2} \Sigma^3 \left[ 1 + O\left(\frac{V}{c}\right) \right] = -\gamma_L \boldsymbol{\omega} \cdot \mathbf{S} \left[ 1 + O\left(\frac{V}{c}\right) \right]. \quad (62)$$

which is the usual “*spin-rotation coupling*” term [6, 7, 8, 12].

#### 4.4 Energy operator in the two frames with Ryder’s rotating tetrad

Since Ryder’s [12] first tetrad ( $u_\alpha$ ), Eq. (33) above, is “adapted” in the sense of Eq. (19) to the inertial frame  $F'$ , it is interesting to compute the energy operator associated in the inertial frame  $F'$  with this tetrad. We checked that, as was found by Ryder, the spin connection matrices (13) for this tetrad field ( $u_\alpha$ ) are

$$\Gamma_0^\# = -\frac{i\omega}{2c} \Sigma^3, \quad \Gamma_j^\# = 0. \quad (63)$$

The tetrad  $(u_\alpha)$  is related to the natural basis  $(\partial'_\mu)$  by Eqs. (34)–(35). We thus transform immediately the connection matrices to the natural basis, getting the same:

$$\Gamma_0 = -\frac{i\omega}{2c}\Sigma^3, \quad \Gamma_j = 0. \quad (64)$$

We get also from (9) and (34)–(35), using then (7):

$$\alpha^0 = \gamma^{\#0}, \quad \alpha^1 = \cos \omega t \alpha^{\#1} - \sin \omega t \alpha^{\#2}, \quad \alpha^2 = \sin \omega t \alpha^{\#1} + \cos \omega t \alpha^{\#2}, \quad \alpha^3 = \alpha^{\#3}. \quad (65)$$

We note that here again  $\gamma^0 = \alpha^0 = \gamma^{\#0}$  is constant, so the Hamiltonian is Hermitian, Eq. (18). From (31), (64), and (65), we find the explicit expression of the energy operator  $E'_3 = H'_3$ , Eq. (6):

$$\begin{aligned} H'_3 &= mc^2 \alpha^0 - i\hbar c \alpha^j \partial'_j - i\hbar c \Gamma_0 \\ &= mc^2 \gamma^{\#0} - i\hbar c [\alpha^{\#1}(\cos \omega t \partial'_x + \sin \omega t \partial'_y) + \alpha^{\#2}(-\sin \omega t \partial'_x + \cos \omega t \partial'_y) + \alpha^{\#3} \partial'_z] \\ &\quad - i\hbar c \Gamma_0, \\ H'_3 &= mc^2 \gamma^{\#0} - i\hbar c \alpha^{\#j} \partial_j - \frac{\hbar \omega}{2} \Sigma^3. \end{aligned} \quad (66)$$

[Recall that  $(\partial_\mu)$  is the natural basis of the *rotating* coordinates.] Thus, with Ryder's tetrad, we find that the DFW energy operator in the *inertial* frame  $F'$  does contain the spin-rotation coupling term  $-\frac{\hbar \omega}{2} \Sigma^3 = -\boldsymbol{\omega} \cdot \mathbf{S}$ . This is certainly unexpected physically. Also, by comparing  $H'_1$  with  $H'_3$  [Eqs. (26) and (66)], we have a clear confirmation of the non-uniqueness of the DFW Hamiltonian and energy operator. The energy operators  $H'_1$  and  $H'_3$  were known in advance to be physically inequivalent [17].

Although Ryder's tetrad is not “adapted” to the rotating frame  $F$  in the sense of Eq. (19), it will turn out to be interesting to have the precise expression of the Hamiltonian and energy operator in that frame  $F$  [in the coordinates  $(x^\mu)$ , Eq. (24)] with this tetrad. That precise expression was not given by Ryder [12], who wrote: “The Dirac equation (4) then, on rearrangement, is found to have a  $\boldsymbol{\sigma} \cdot \boldsymbol{\omega}$  ( $= \omega \sigma^3$  here) contribution to the Hamiltonian—a spin-rotation coupling term exactly as predicted by Mashhoon.” From the expression (33) of that tetrad as function of the natural basis of the coordinates  $(x^\mu)$ , and from (63), we get once again for the connection matrices [this time in the coordinates  $(x^\mu)$ ]:

$$\Gamma_0 = -\frac{i\omega}{2c}\Sigma^3, \quad \Gamma_j = 0, \quad (67)$$

and we get the  $\gamma^\mu$  matrices (9),

$$\gamma^0 = \gamma^{\#0}, \quad \gamma^1 = \frac{\omega y}{c} \gamma^{\#0} + \gamma^{\#1}, \quad \gamma^2 = -\frac{\omega x}{c} \gamma^{\#0} + \gamma^{\#2}, \quad \gamma^3 = \gamma^{\#3} \quad (68)$$

[thus once more the hermiticity condition (18) is verified], whence for the  $\alpha^\mu$  matrices in Eq. (7):

$$\alpha^0 = \gamma^{\#0}, \quad \alpha^1 = \frac{\omega y}{c} \mathbf{1}_4 + \alpha^{\#1}, \quad \alpha^2 = -\frac{\omega x}{c} \mathbf{1}_4 + \alpha^{\#2}, \quad \alpha^3 = \alpha^{\#3}. \quad (69)$$

Therefore, the energy operator  $E_3 = H_3$ , Eq. (6), is now:

$$\begin{aligned} H_3 &= mc^2 \alpha^0 - i\hbar c \alpha^j \partial_j - i\hbar c \Gamma_0 \\ &= mc^2 \gamma^{\#0} - i\hbar c \left[ \alpha^{\#j} \partial_j + \frac{\omega}{c} (y \partial_x - x \partial_y) + \Gamma_0 \right], \end{aligned} \quad (70)$$

$$H_3 = H'_3 - \boldsymbol{\omega} \cdot \mathbf{L}. \quad (71)$$

Remembering Eq. (67), we see that with Ryder's (first) tetrad, the energy operator in the rotating frame F has indeed the spin-rotation coupling term  $-\frac{\hbar\omega}{2}\Sigma^3 = -\boldsymbol{\omega} \cdot \mathbf{S}$ —as has the energy operator with this tetrad but in the inertial frame F', Eq. (66). Also, these two energy operators differ from one another only by the *angular momentum* term—just as we found also with the Cartesian tetrad, Eqs. (26) and (32).

## 4.5 The general relation between the Hamiltonians in two frames in relative rotation

It turns out to be a general fact that the Dirac Hamiltonian operators in two reference frames in relative rotation differ only by the angular momentum term, if they correspond to the same tetrad field. In a general Lorentzian spacetime  $(V, \mathbf{g})$ , consider a general reference frame R', defined by a chart  $\chi' : X \mapsto (x'^\mu) = (ct', x', y', z')$ , and define another reference frame R by a chart  $\chi$  deduced from  $\chi'$  by a transformation generalizing (24):

$$t = t', \quad x = x' \cos \phi(t) + y' \sin \phi(t), \quad y = -x' \sin \phi(t) + y' \cos \phi(t), \quad z = z'. \quad (72)$$

So the spatial coordinate vector  $\mathbf{r} \equiv (x, y, z)$ , at least, is undergoing a rotation, at a variable rate  $\omega \equiv \dot{\phi} \equiv d\phi/dt$ , with respect to the space browsed by the coordinates  $(x'^j)$ . This corresponds to a rotation in physical space if V is endowed with the Minkowski metric  $\gamma$  (with possibly  $\gamma = \mathbf{g}$  as a particular

case) and the chart  $\chi'$  is Cartesian for  $\gamma$ . The Dirac Hamiltonian (6) rewrites immediately, in the most general case, as

$$H = i \frac{\partial}{\partial t} + \frac{m}{g^{00}} \gamma^0 - i \frac{\gamma^0 \gamma^\mu D_\mu}{g^{00}} \quad (\hbar = 1 = c). \quad (73)$$

On a general coordinate change,  $\gamma^\mu D_\mu$  is invariant (for a given tetrad field, of course), due to the transformation behaviours of  $\gamma^\mu$ ,  $\Gamma_\mu$ , and  $\partial_\mu$ . On the coordinate change (72),  $\gamma^0$  and  $g^{00}$  are invariant. Therefore, we have

$$H' - H = i \left( \frac{\partial}{\partial t'} - \frac{\partial}{\partial t} \right) = i \frac{\partial x^j}{\partial t'} \frac{\partial}{\partial x^j} = i \dot{\phi} \left( y \frac{\partial}{\partial x} - x \frac{\partial}{\partial y} \right), \quad (74)$$

that is,

$$H - H' = -\boldsymbol{\omega} \cdot \mathbf{L}, \quad \boldsymbol{\omega} \equiv (\omega, 0, 0), \quad \mathbf{L} \equiv \mathbf{r} \wedge (-i\hbar \nabla), \quad (75)$$

as announced.<sup>8</sup>

## 5 Conclusion

The predictions of the spin-rotation coupling term for a particle obeying the covariant Dirac equation (5) have considered a tetrad field which is undergoing more or less the same rotation as the rotating reference frame itself [7, 8, 12]. As suggested by Hehl & Ni [7] and by Ryder [12], to make this precise one should use the notion of the Fermi-Walker transport or derivative. By using the Fermi-Walker derivative one may indeed define rigorously the rotation rate of the spatial triad associated with an orthonormal tetrad, for a general reference frame in a general spacetime [17]. This rotation rate is the spatial tensor field  $\Xi$  in Eq. (22). That definition needs that one considers an “adapted” tetrad to the reference frame considered, i.e., one such that the time-like vector of the tetrad is the four-velocity of the reference frame, Eq. (19). The rotation rate of the reference frame should be precisely defined also as a spatial tensor field  $\Omega$  and can indeed be, Eq. (23).

For a uniformly rotating frame in the Minkowski spacetime, we succeeded at defining an *adapted* tetrad field which verifies  $\Xi = \Omega$  almost exactly. With this tetrad field, the energy operator in the rotating frame does have the

---

<sup>8</sup> In particular, the Hamiltonian in the inertial frame  $F'$  and with the adapted tetrad (42)-(43) is:  $H'_2 = H_2 + \boldsymbol{\omega} \cdot \mathbf{L}$ , with  $H_2$  given by Eq. (61). This is also the energy operator.

spin-rotation coupling term, Eq. (62). We also wrote explicitly the energy operator with Ryder’s rotating tetrad field [12], which does involve this term, too—although Ryder’s tetrad is adapted to the inertial frame, not to the rotating frame.

However, the three tetrad fields investigated in the present work provide three different Hamiltonians in the inertial frame, as well as three different Hamiltonians in the rotating frame. (In each case, the Hamiltonian coincides with the energy operator.) Moreover, those tetrads that provide the spin-rotation coupling term in the energy operator of the rotating frame, give it also in the energy operator of the *inertial frame*. In fact, we find fully generally that the Hamiltonian operators in two reference frames in relative rotation, but corresponding to the same tetrad field, differ *only* by the *angular momentum* term, Eq. (75). Thus, if the Hamiltonian involves spin-rotation coupling in the rotating frame, and if one keeps the same tetrad, then the corresponding Hamiltonian in the inertial frame *must* also involve spin-rotation coupling, which is certainly unexpected physically. Therefore, if the spin-rotation coupling is to exist for a Dirac particle, it means that two different tetrad fields must be chosen for two different reference frames. Thus, for each given reference frame, a tetrad field adapted to that reference frame should be chosen. Then, to get the relevant rotation rate in the spin-rotation coupling term, one has to impose that the rotation rate of the triad is indeed that of the reference frame:  $\Xi = \Omega$ . That is, if the spin-rotation coupling is to exist for a Dirac particle, Variant *a*) of Framework I is the correct scheme to select the tetrad field. As we saw, this is difficult to implement already for the simple case of a uniform rotation in a Minkowski spacetime—not to speak of a general situation.

One may consider that the choice of a tetrad field should be valid for any reference frame instead. Framework II is the only currently available one that ensures this while providing unambiguous Dirac Hamiltonian and energy operators. It assumes that the metric can be put in the form (3) in some chart: preferably a global one of course, in which case, by setting  $\gamma_{\mu\nu} \equiv \eta_{\mu\nu}$  in that chart, one endows the spacetime with the Minkowski metric  $\gamma$ , related simply to the physical metric  $\mathbf{g}$ . In the case of a Minkowski spacetime ( $\mathbf{g} = \gamma$ ), this framework leads to selecting any “Cartesian tetrad”. It predicts no spin-rotation coupling. Thus, experiments should decide.

## References

- [1] S. A. Werner, J. L. Staudenmann, and R. Colella, “Effect of Earth’s rotation on the quantum mechanical phase of the neutron,” *Phys. Rev. Lett.* **42**, 1103–1106 (1979).
- [2] M. Arminjon, “Main effects of the Earth’s rotation on the stationary states of ultra-cold neutrons,” *Phys. Lett. A* **372**, 2196–2200 (2008). [See also arXiv:0708.3204v2 (quant-ph)]
- [3] J. Kuroiwa, M. Kasai, and T. Futamase, “A treatment of general relativistic effects in quantum interference,” *Phys. Lett. A* **182**, 330–334 (1993).
- [4] V. S. Morozova and B. J. Ahmedov, “Quantum interference effects in slowly rotating NUT space-time,” *Int. J. Mod. Phys. D* **18**, 107–118 (2009). [arXiv:0804.2786v2 (gr-qc)]
- [5] B. Mashhoon, “On the coupling of intrinsic spin with the rotation of the earth,” *Phys. Lett. A* **198**, 9–13 (1995).
- [6] B. Mashhoon, “Neutron interferometry in a rotating frame of reference,” *Phys. Rev. Lett.* **61**, 2639–2642 (1988).
- [7] F. W. Hehl and W. T. Ni, “Inertial effects of a Dirac particle,” *Phys. Rev. D* **42**, 2045–2048 (1990).
- [8] Y. Q. Cai and G. Papini, “Neutrino helicity flip from gravity-spin coupling,” *Phys. Rev. Lett.* **66**, 1259–1262 (1991).
- [9] D. R. Brill and J. A. Wheeler, “Interaction of neutrinos and gravitational fields,” *Rev. Modern Phys.* **29**, 465–479 (1957). Erratum: *Rev. Modern Phys.* **33**, 623–624 (1961).
- [10] T. C. Chapman and D. J. Leiter, “On the generally covariant Dirac equation,” *Am. J. Phys.* **44**, No. 9, 858–862 (1976).
- [11] C. J. Isham, “Spinor fields in four dimensional space-time,” *Proc. Roy. Soc. London A* **364**, 591–599 (1978).
- [12] L. Ryder, “Spin-rotation coupling and Fermi-Walker transport,” *Gen. Relativ. Gravit.* **40**, 1111–1115 (2008).

- [13] M. Arminjon and F. Reifler, “A non-uniqueness problem of the Dirac theory in a curved spacetime,” *Ann. Phys. (Berlin)* **523**, 531–551 (2011). [arXiv:0905.3686 (gr-qc)]
- [14] M. V. Gorbatenko and V. P. Neznamov, “Solution of the problem of uniqueness and hermiticity of Hamiltonians for Dirac particles in gravitational fields,” *Phys. Rev. D* **82**, 104056 (2010). [arXiv:1007.4631v1 (gr-qc)]
- [15] M. Leclerc, “Hermitian Dirac Hamiltonian in the time-dependent gravitational field,” *Class. Quant. Grav.* **23**, 4013–4020 (2006). [arXiv:gr-qc/0511060v3]
- [16] M. Arminjon, “A simpler solution of the non-uniqueness problem of the Dirac theory,” submitted for publication. [arXiv:1205.3386v2 (math-ph)]
- [17] M. Arminjon, “A solution of the non-uniqueness problem of the Dirac Hamiltonian and energy operators,” *Ann. Phys. (Berlin)* **523**, 1008–1028 (2011). [Pre-peer-review version: arXiv:1107.4556v2 (gr-qc)].
- [18] C. Cattaneo, “General relativity: relative standard mass, momentum, energy and gravitational field in a general system of reference,” *il Nuovo Cimento* **10**, 318–337 (1958).
- [19] J. von Weyssenhof, “Metrisches Feld und Gravitationsfeld,” *Bull. Acad. Polon. Sci., Sect. A* **252** (1937). (Quoted by Cattaneo [18].)
- [20] M. Arminjon and F. Reifler, “Basic quantum mechanics for three Dirac equations in a curved spacetime,” *Braz. J. Phys.* **40**, 242–255 (2010). [arXiv:0807.0570 (gr-qc)].
- [21] M. Arminjon and F. Reifler, “Four-vector vs. four-scalar representation of the Dirac wave function,” *Int. J. Geom. Meth. Mod. Phys.* **9**, No. 4, 1250026 (2012). [arXiv:1012.2327v2 (gr-qc)]
- [22] S. S. Chern, W. H. Chen, and K. S. Lam, *Lectures on Differential Geometry* (Singapore: World Scientific 1999), pp. 113–121.
- [23] W. Pauli, “Contributions mathématiques à la théorie des matrices de Dirac,” *Ann. Inst. Henri Poincaré* **6**, 109–136 (1936).
- [24] L. Parker, “One-electron atom as a probe of spacetime curvature,” *Phys. Rev. D* **22**, 1922–1934 (1980).

- [25] X. Huang and L. Parker, “Hermiticity of the Dirac Hamiltonian in curved spacetime,” *Phys. Rev. D* **79**, 024020 (2009). [arXiv:0811.2296 (gr-qc)]
- [26] B. Mashhoon and U. Muench, “Length measurement in accelerated systems,” *Ann. Phys. (Berlin)* **11**, 532–547 (2002). [arXiv:gr-qc/0206082v1]
- [27] J. W. Maluf, F. F. Faria, and S. C. Ulhoa, “On reference frames in spacetime and gravitational energy in freely falling frames,” *Class. Quant. Grav.* **24**, 2743–2754 (2007). [arXiv:0704.0986v1 (gr-qc)]
- [28] M. Arminjon, “Space isotropy and weak equivalence principle in a scalar theory of gravity,” *Braz. J. Phys.* **36**, 177–189 (2006). [arXiv:gr-qc/0412085]
- [29] R. T. Jantzen, P. Carini, and D. Bini, “The many faces of gravitoelectromagnetism,” *Ann. Phys. (New York)* **215**, 1–50 (1992). [arXiv:gr-qc/0106043]